

Cosmic web analysis in the SDSS main galaxy sample and implications for galaxy colors

Florent Leclercq

Institute of Cosmology and Gravitation, University of Portsmouth

<http://icg.port.ac.uk/~leclercq/>



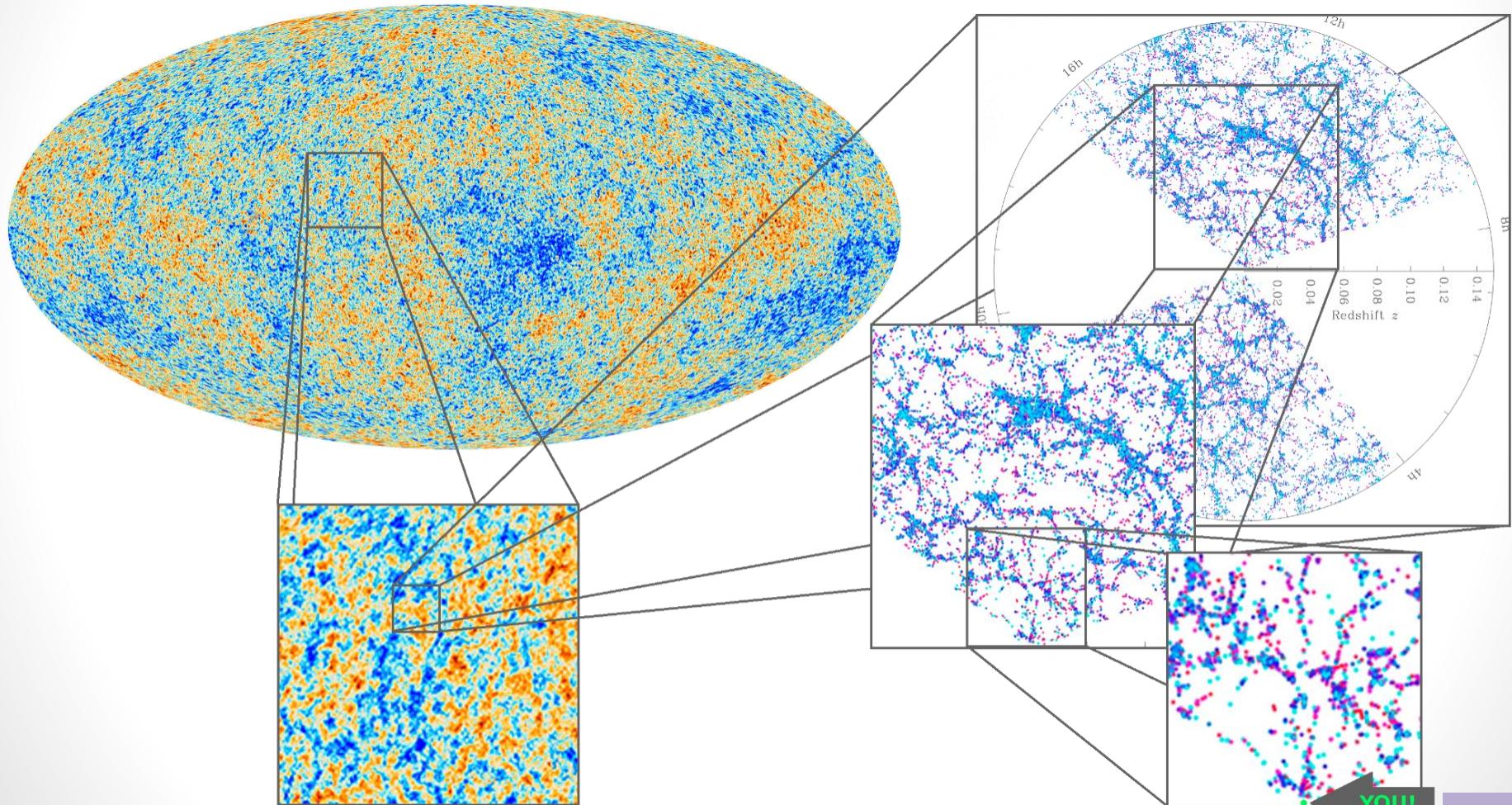
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In collaboration with:

Jens Jasche (ExC Universe, Garching), Guilhem Lavaux (IAP),
Will Percival (ICG), Benjamin Wandelt (IAP/U. Illinois)

The big picture: the Universe is highly structured

You are here. Make the best of it...



Planck collaboration (2013-2015)

M. Blanton and the Sloan Digital Sky Survey (2010-2013)

What we want to know from the LSS

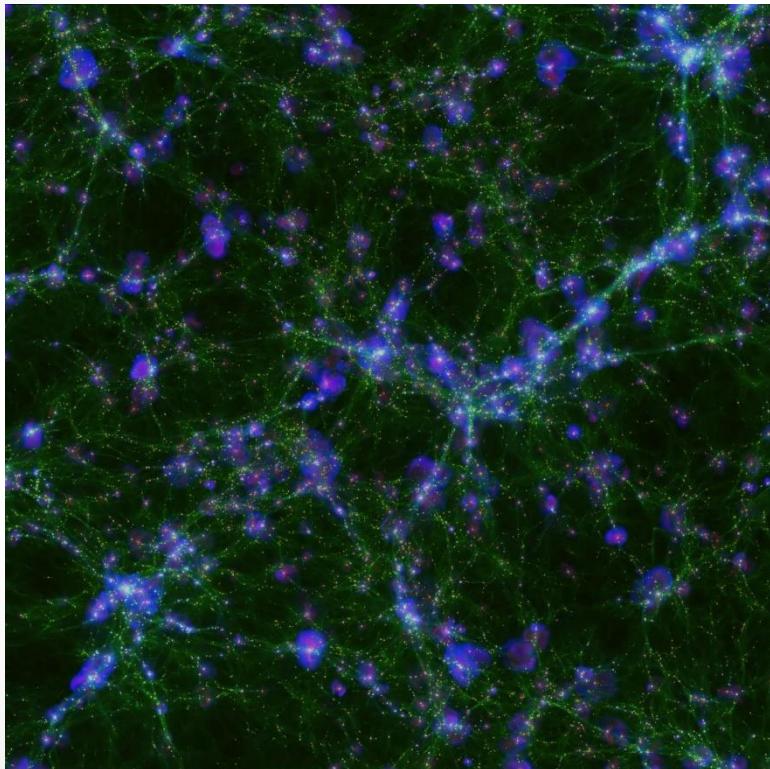
The LSS is a vast source of knowledge:

- **Cosmology:**

- Cosmological parameters and tests of Λ CDM,
- Physical nature of the dark components,
- Geometry of the Universe,
- Tests of General Relativity,
- Initial conditions and link to high energy physics

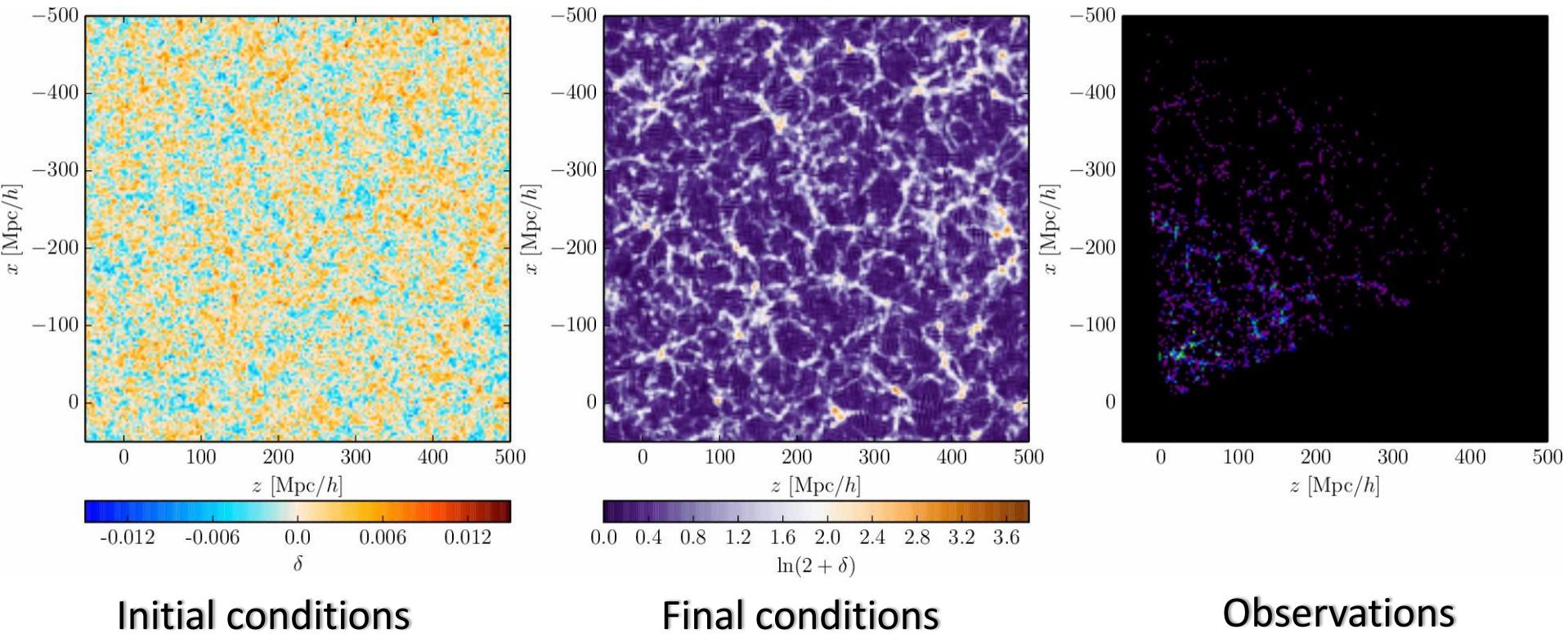
- **Astrophysics:** galaxy formation and evolution as a function of their environment

- Galaxy properties (colors, chemical composition, shapes),
- Intrinsic alignments



Y. Dubois (PI), Horizon AGN simulation (2014-2016)

The BORG SDSS run



Initial conditions

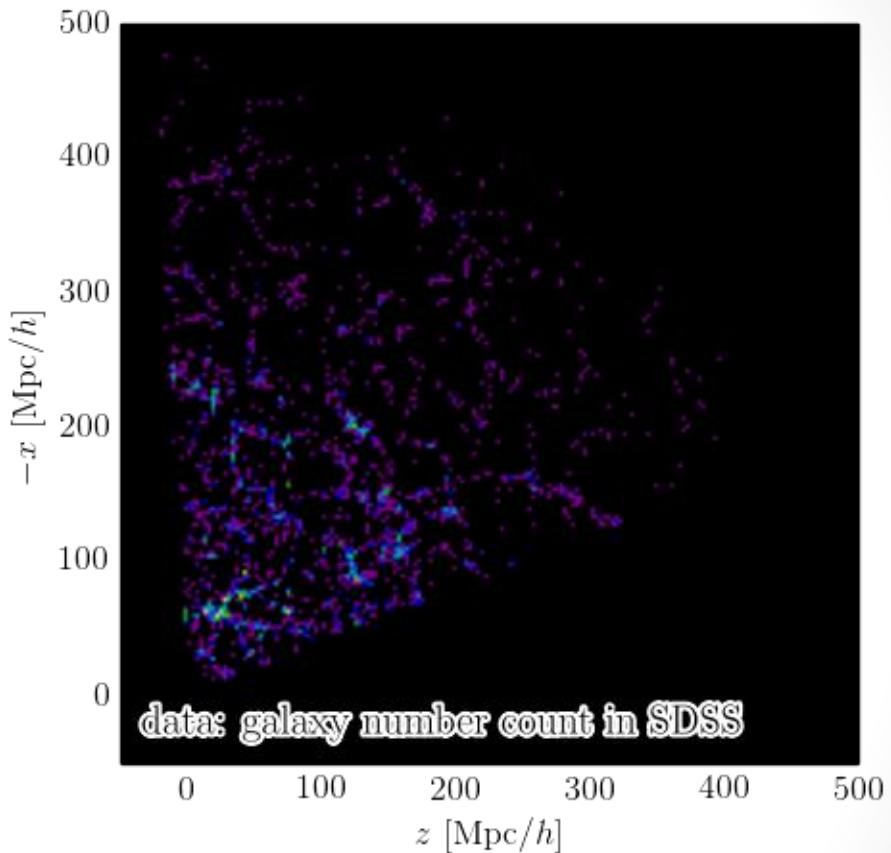
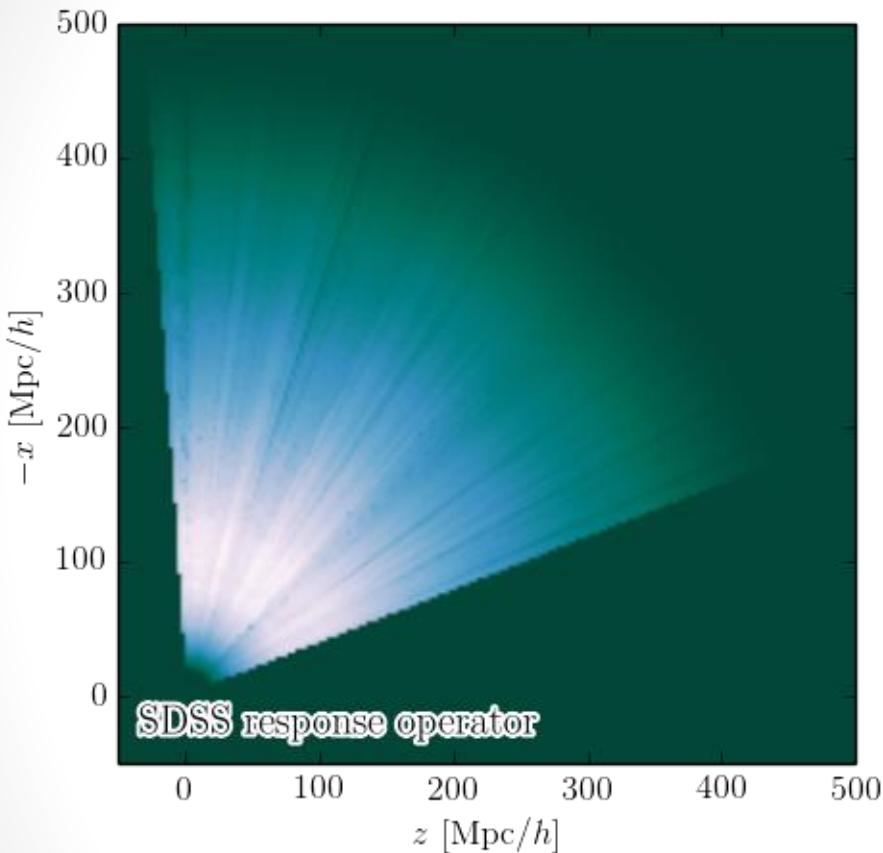
Final conditions

Observations

334,074 galaxies, \approx 17 millions parameters, 3 TB of primary data products,
12,000 samples, \approx 250,000 data model evaluations, 10 months on 32 cores

Jasche, FL & Wandelt 2015, arXiv:1409.6308

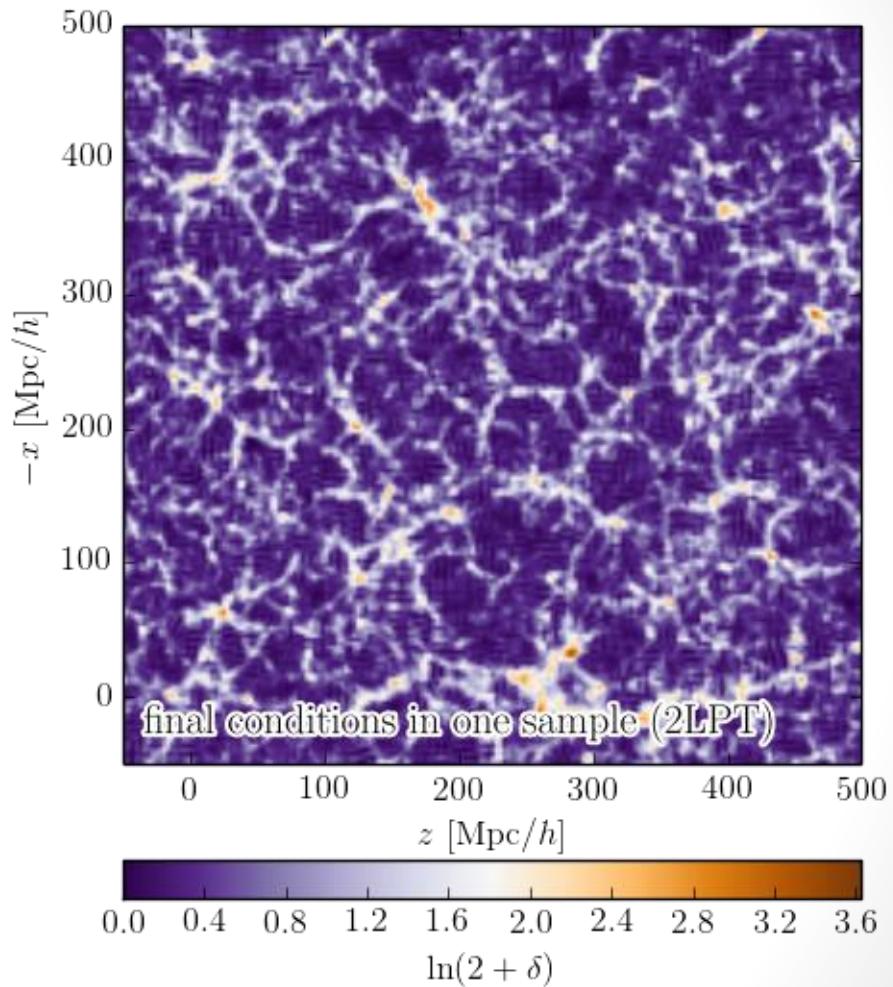
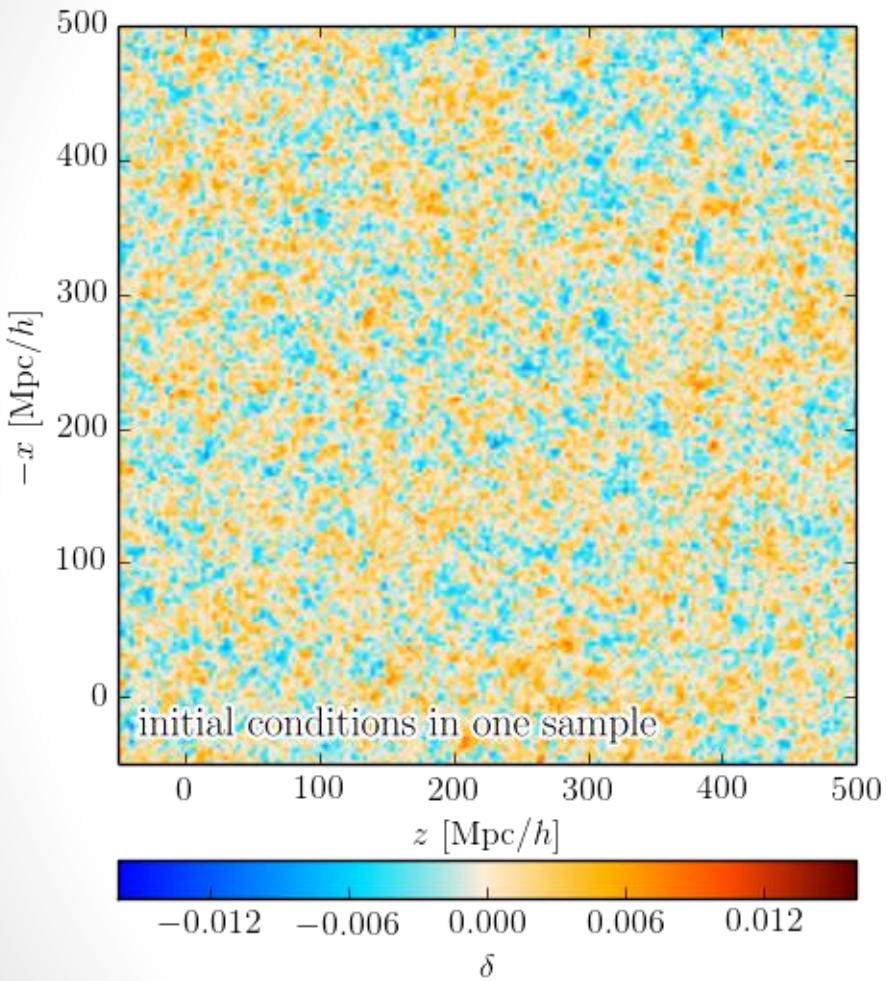
Bayesian chrono-cosmography from SDSS DR7



Data

Jasche, FL & Wandelt 2015, arXiv:1409.6308

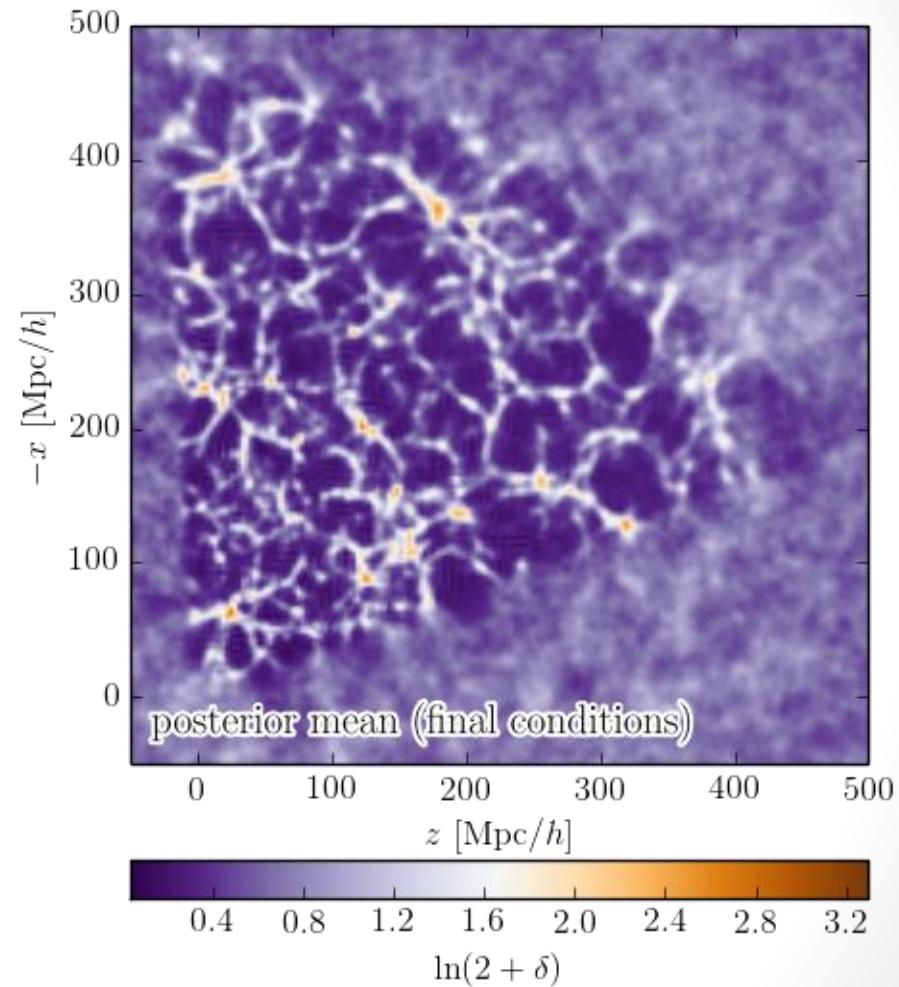
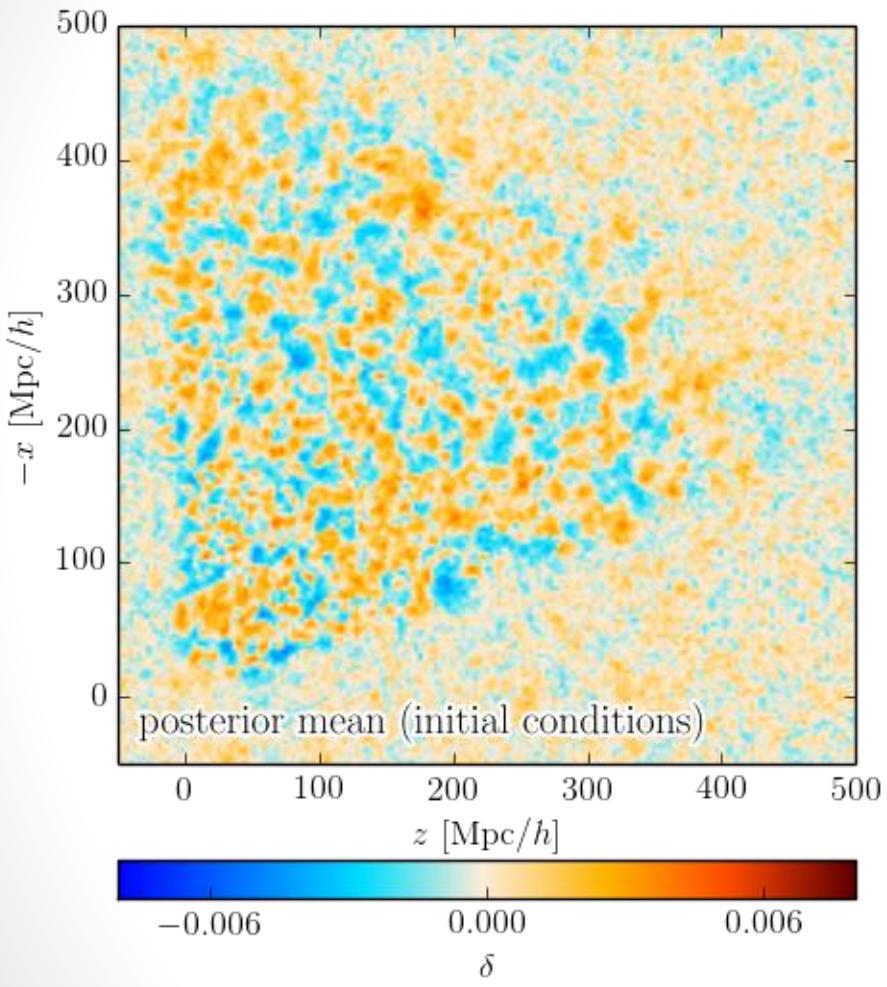
Bayesian chrono-cosmography from SDSS DR7



One sample

Jasche, FL & Wandelt 2015, arXiv:1409.6308

Bayesian chrono-cosmography from SDSS DR7

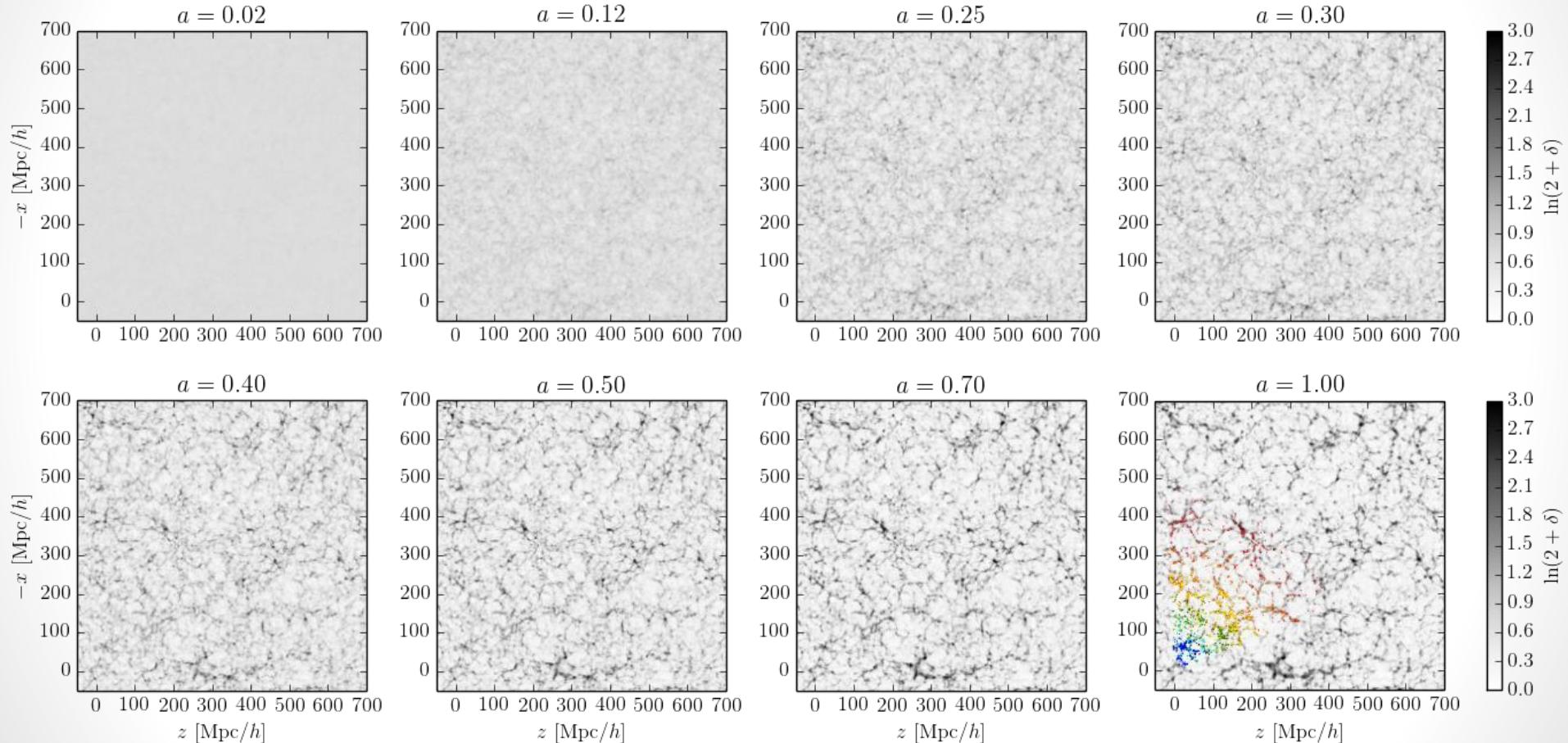


Posterior mean

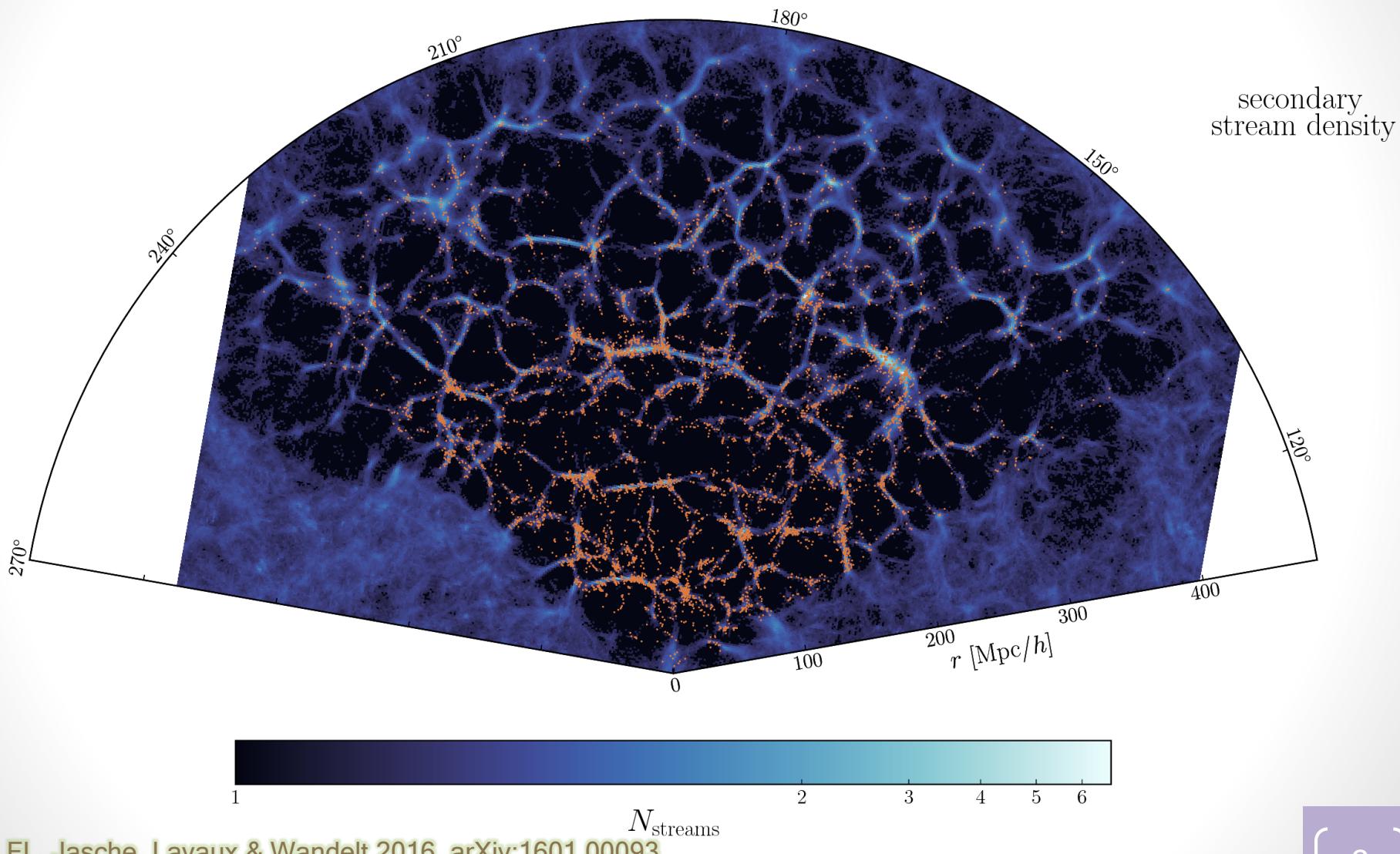
Jasche, FL & Wandelt 2015, arXiv:1409.6308

(7)

Evolution of cosmic structure

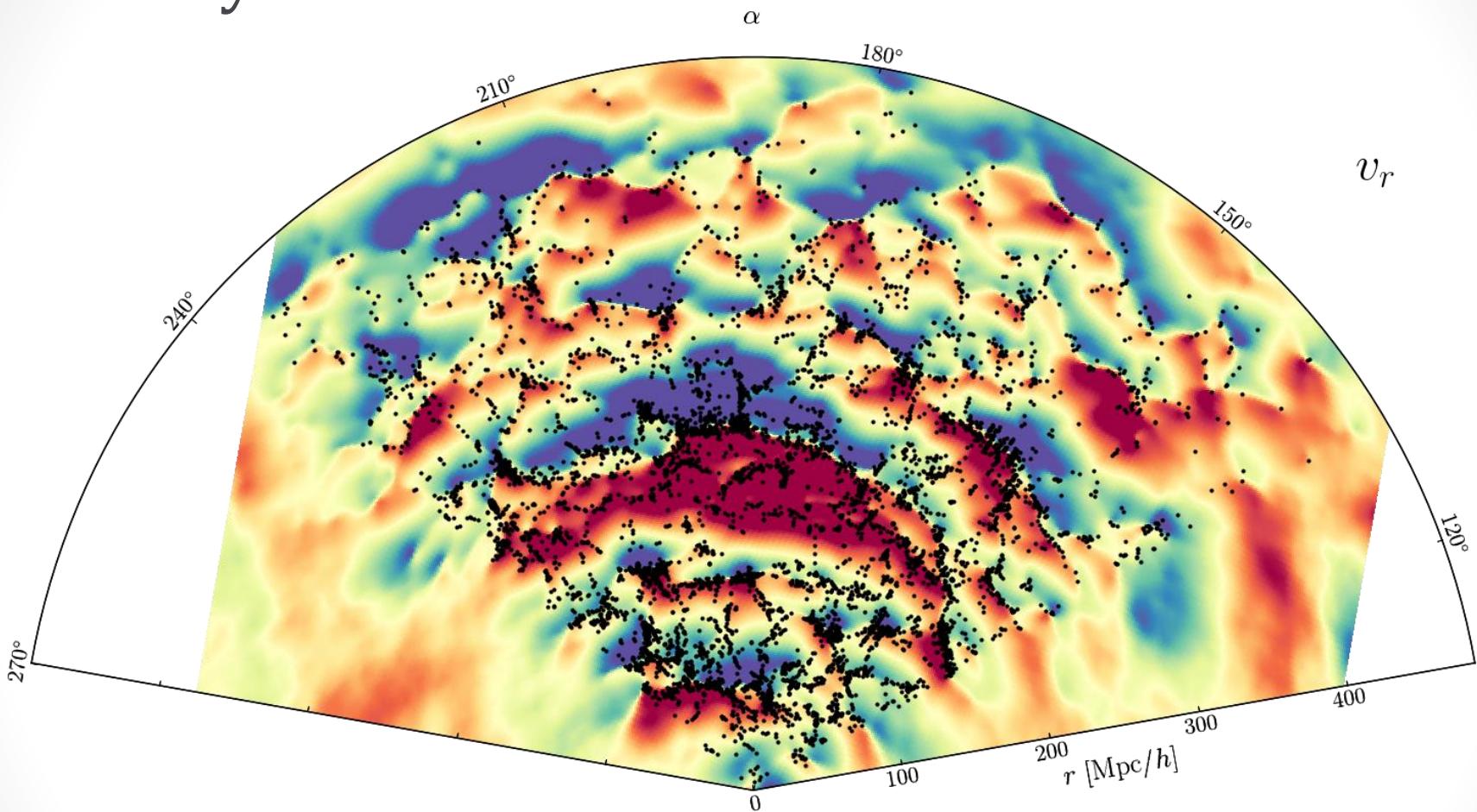


Dark matter stream density



FL, Jasche, Lavaux & Wandelt 2016, arXiv:1601.00093

Velocity field

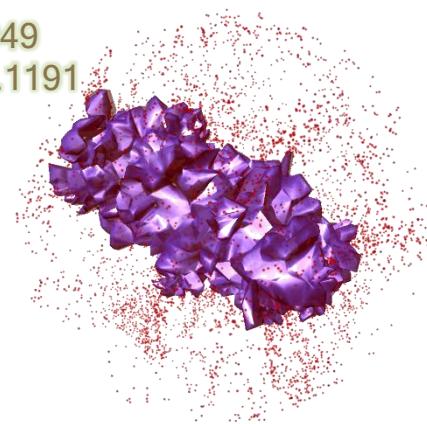


Cosmic web elements: some algorithms

- “**Structure finders**” focus on one element at a time

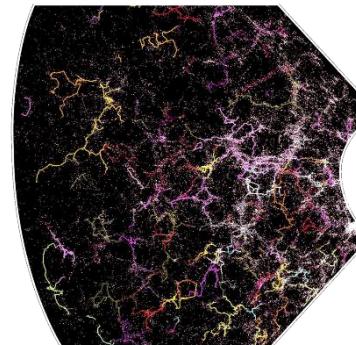
- **ZOBOV/VIDE**

Neyrinck 2008, arXiv:0712.3049
Sutter *et al.* 2015, arXiv:1406.1191



- **DisPerSE**

Sousbie 2011, arXiv:1009.4015
Sousbie *et al.* 2011, arXiv:1009.4014



- “**Classifiers**” dissect the cosmic web all at once

- The **T-web** (tidal field tensor)

Hahn *et al.* 2007, arXiv:astro-ph/0610280

- **DIVA** (Lagrangian displacement field, potential structures)

Lavaux & Wandelt 2010, arXiv:0906.4101

- **ORIGAMI** (particle crossings)

Falck, Neyrinck & Szalay 2012, arXiv:1201.2353

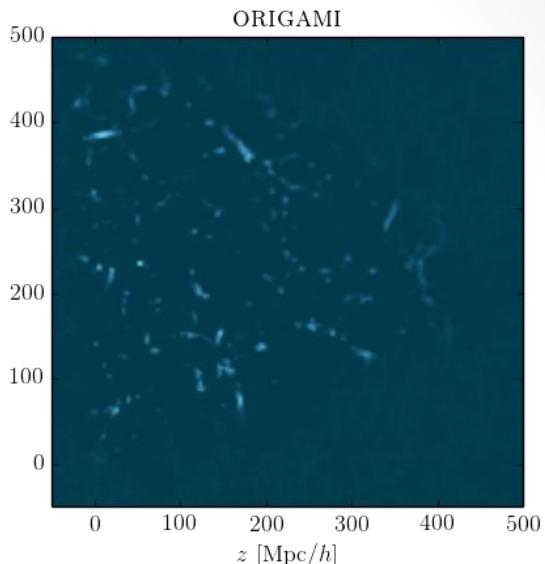
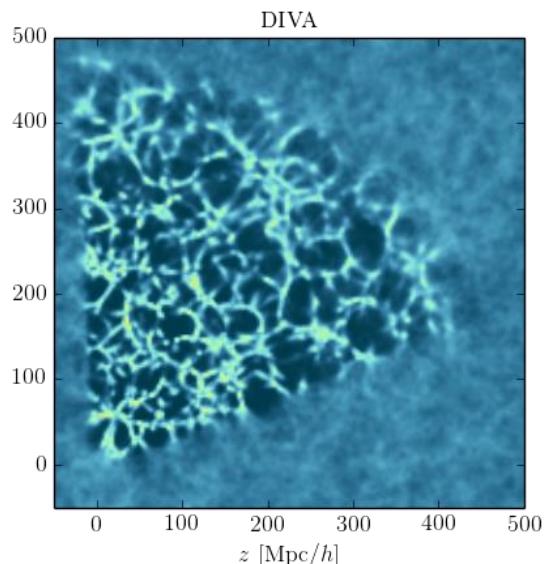
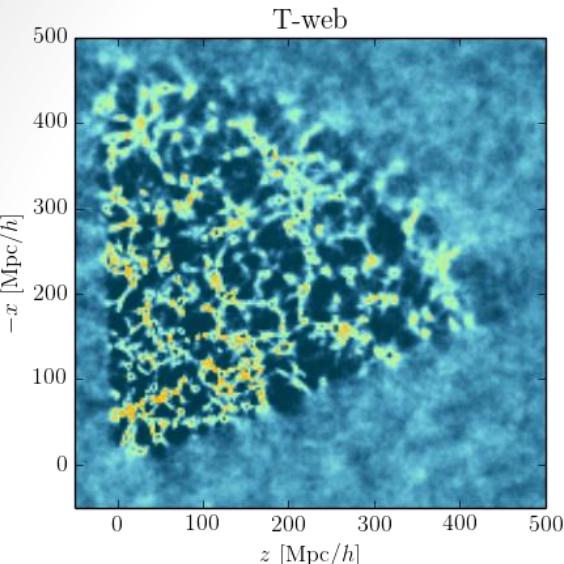
- **LICH** (Lagrangian displacement field, potential and vortical structures)

FL, Jasche, Lavaux & Wandelt 2016, arXiv:1601.00093

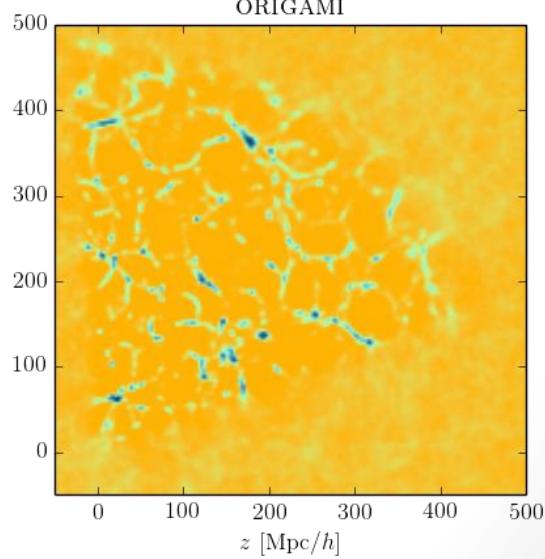
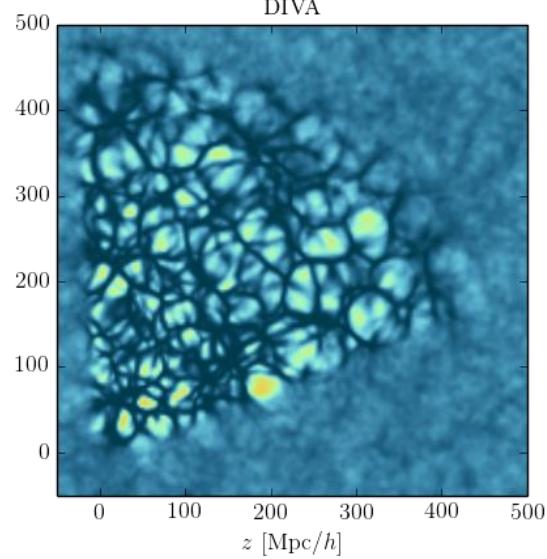
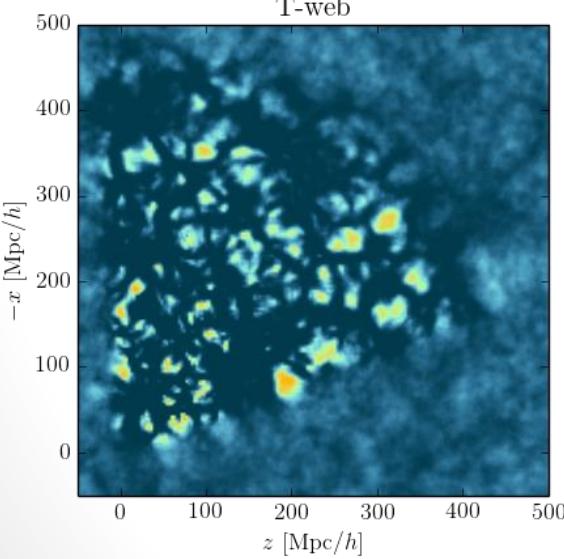
and many others...

Comparing classifiers

Filaments



Voids



FL, Jasche & Wandelt 2015a, arXiv:1502.02690

FL, Jasche, Lavaux & Wandelt 2016, arXiv:1601.00093

Which is the best classifier?

- The framework is provided by **Bayesian decision theory** and **information theory**
- The idea is to maximize a utility function

$$U(\xi) = \langle U(d, T, \xi) \rangle_{\mathcal{P}(d, T | \xi)}$$

- An important notion: the **mutual information** between two random variables (or **expected information gain**)

$$\begin{aligned} I[X : Y] &\equiv D_{\text{KL}}[\mathcal{P}(x, y) || \mathcal{P}(x)\mathcal{P}(y)] \\ &= \sum_{x \in \mathcal{X}, y \in \mathcal{Y}} \mathcal{P}(x, y) \log_2 \left(\frac{\mathcal{P}(x, y)}{\mathcal{P}(x)\mathcal{P}(y)} \right) \end{aligned}$$

1. Utility for parameter inference:

example: cosmic web analysis

- **Example:** Which classifier produces the most “surprising” cosmic web maps when looking at the data?

$$U_1(\xi) = I[T : d | \xi]$$

classification data

2. Utility for model selection:

example: dark energy equation of state

- **Example:** Let us consider three dark energy models with

$$w = -0.9, w = -1, w = -1.1$$

Which classifier separates them better?

$$U_2(\xi) = I[\mathcal{M} : \mathcal{R}(d) | \xi]$$

model classifier mixture distribution

$$\mathcal{R}(d) \equiv \frac{\mathcal{P}(T(\vec{x}_k) | d, \mathcal{M}_1) + \mathcal{P}(T(\vec{x}_k) | d, \mathcal{M}_2)}{2}$$

3. Utility for prediction of new data:

example: galaxy colors

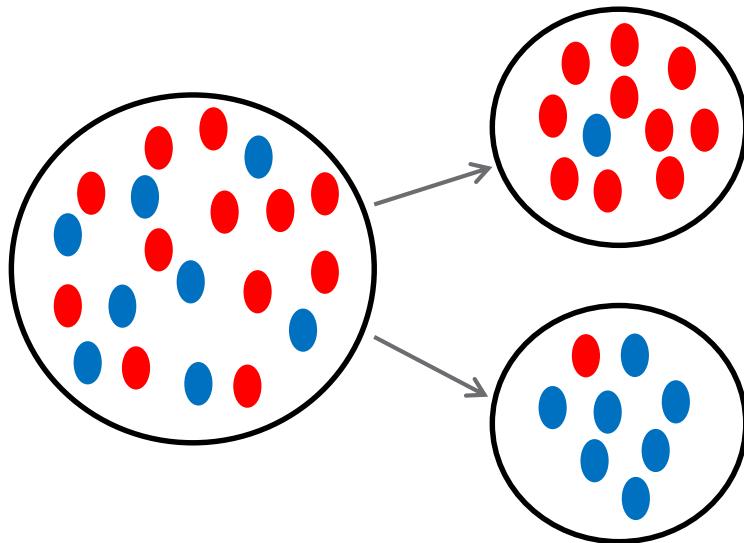
- **Example:** *So far we have not used galaxy colors. Which classifier predicts them best?*
- Maximize the **expected information gain** on some new quantity

$$U_3(\xi) = I[c:T|\xi]$$

predicted data classification

3. Utility for prediction of new data: example: galaxy colors

- How to compute the information gain?



parent entropy:

$$H = -\frac{8}{20} \log_2 \left(\frac{8}{20} \right) - \frac{12}{20} \log_2 \left(\frac{12}{20} \right) = 0.9709$$

child1 entropy:

$$H = -\frac{10}{11} \log_2 \left(\frac{10}{11} \right) - \frac{1}{11} \log_2 \left(\frac{1}{11} \right) = 0.4395$$

child2 entropy:

$$H = -\frac{8}{9} \log_2 \left(\frac{8}{9} \right) - \frac{1}{9} \log_2 \left(\frac{1}{9} \right) = 0.5033$$

weighted average entropy of children:

$$\frac{11}{20} \times 0.4395 + \frac{9}{20} \times 0.5033 = 0.4682$$

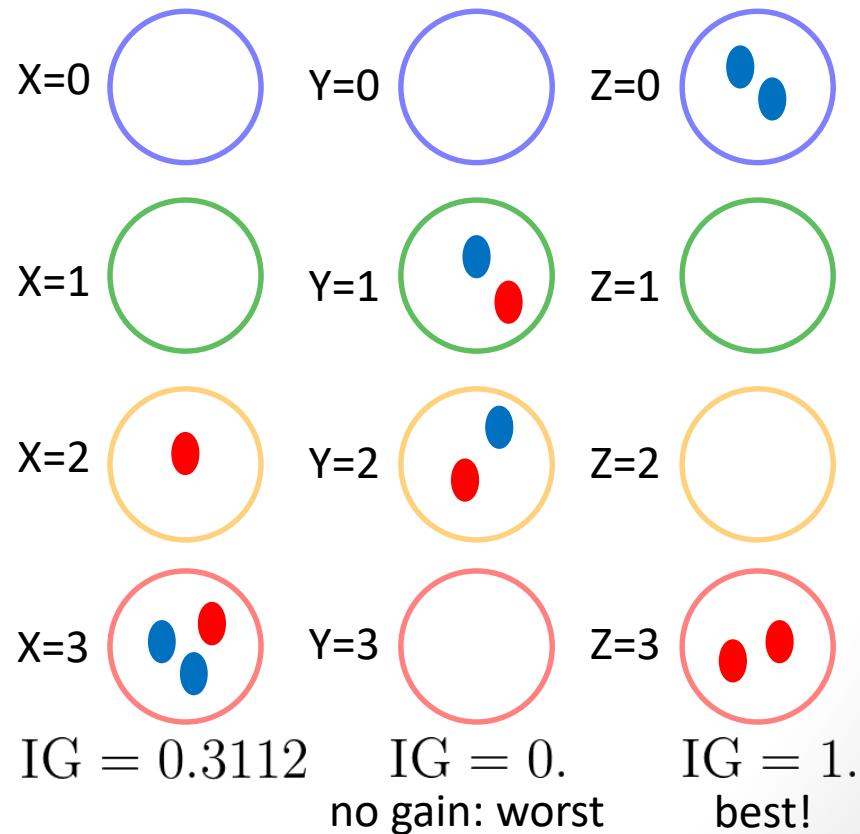
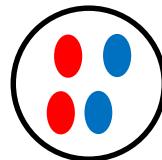
information gain for this split: $0.9709 - 0.4682 = 0.5027$ Sh

3. Utility for prediction of new data:

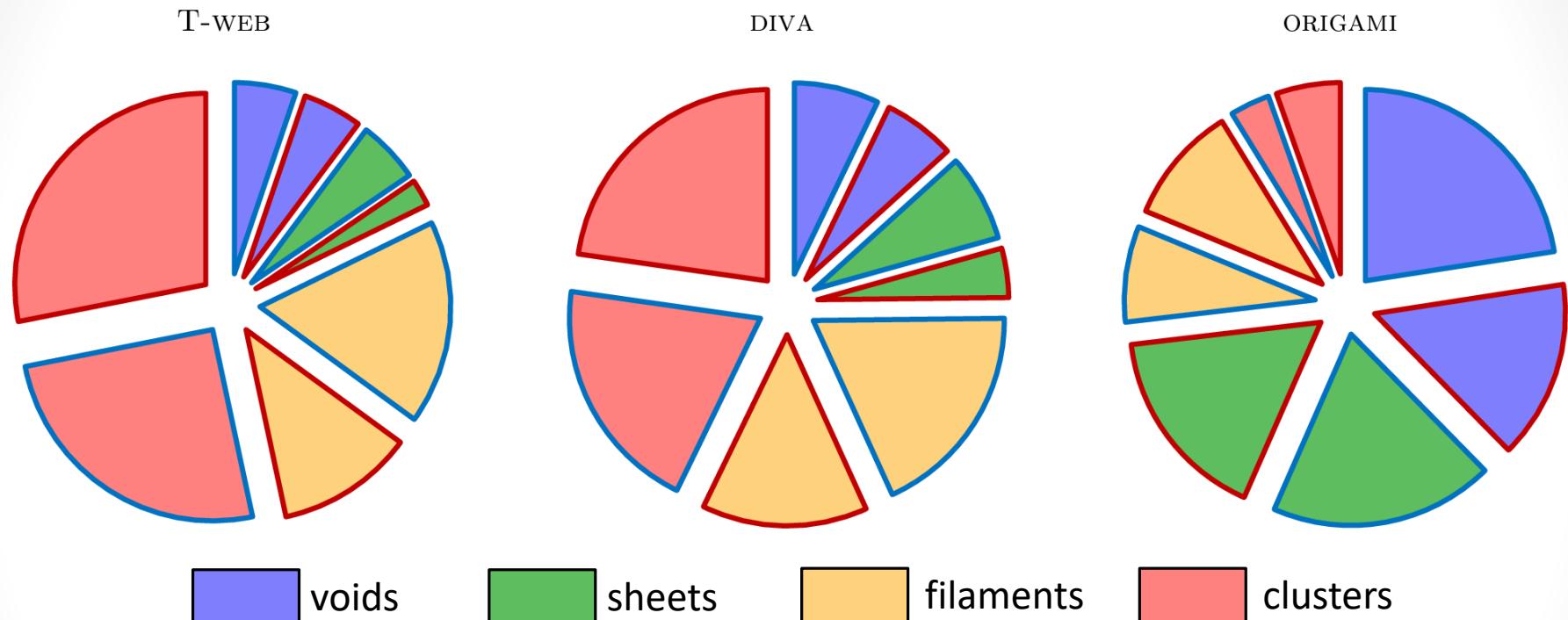
example: galaxy colors

- A **supervised machine learning** problem!
 - 3 **features** = classifications ($X=T\text{-web}$, $Y=\text{DIVA}$, $Z=\text{ORIGAMI}$)
 - 4 **possible values** (0=void, 1=sheet, 2=filament, 3=cluster)
 - 2 **classes** (I=blue, II=red)

X	Y	Z	C
3	2	0	I
3	1	0	I
2	2	3	II
3	1	3	II



Application to SDSS main sample galaxies (367,157 galaxies)



	T-web	DIVA	ORIGAMI
Utility [Sh]	0.0152	0.0101	0.0143

Summary & Conclusions

- **BORG** allows a rich description of the large-scale structure of the Universe (density field, velocity field, stream density...)
- The **cosmic web** can be described using various classifiers (T-web, DIVA, ORIGAMI...)
- **Decision theory** and **information theory** offer a framework to rank classifiers, with utility functions depending on the desired use
- Potential **applications**: galaxy properties, intrinsic alignments...

All maps, catalogs & scripts are publicly available at <http://icg.port.ac.uk/~leclercq/>

References

Jasche & Wandelt 2013, arXiv:1203.3639	(BORG proof of concept)
Jasche, FL & Wandelt 2015, arXiv:1409.6308	(BORG SDSS analysis)
FL, Jasche & Wandelt 2015a, arXiv:1502.02690	(T-web, entropy, relative entropy)
FL, Jasche, Lavaux & Wandelt 2016, arXiv:1601.00093	(DIVA, ORIGAMI, phase-space properties)
FL, Lavaux, Jasche & Wandelt 2016, arXiv:1606.06758	(mutual information, classifier utilities)